Kolya: an open-source package for inclusive semileptonic B decays

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Abstract

We introduce the code kolya, an open-source tool for phenomenological analyses of inclusive semileptonic B meson decays. It contains a library to compute predictions for the total rate and various kinematic moments within the framework of the heavy quark expansion, utilizing the so-called kinetic scheme. The library currently includes power corrections up to $1/m_b^5$. All available QCD perturbative corrections are implemented via interpolation grids for fast numerical evaluation. We also include effects from new physics parameterised as Wilson coefficients of dimension-six operators in the weak effective theory below the electroweak scale. The library is interfaced to CRunDec for easy evaluation of the quark masses and strong coupling constant at different renormalization scales. The library is developed in Python and does not require compilation. It can be used in an interactive Jupyter notebook session.

1 Introduction

Measurements of semileptonic B decays lie at the core of the Belle II and LHCb physics program in the upcoming years. Thanks to relatively large rates and clean experimental signatures, inclusive and exclusive semileptonic decays with a $b \to cl\bar{\nu}_l$ transition ($\ell = e, \mu$) offer a clean avenue for the determinations of $|V_{cb}|$, the element of the Cabibbo-Kobayashi-Maskawa matrix (CKM) which parameterizes the strength of the weak interaction among bottom and charm quarks in the Standard Model of particle physics.

Inclusive determinations of $|V_{cb}|$ exploit that the semileptonic rate $\Gamma_{\rm sl}$ and moments of kinematic spectra can be described with good precision using the heavy quark expansion (HQE) [1–3]. In the HQE, these observables are expressed as a series of non-perturbative HQE elements proportional to increasing powers of the inverse bottom quark mass times the QCD scale parameter, $\Lambda_{\rm QCD}/m_b$. In addition, each order in the HQE also receives corrections expressed as a series expansion in the strong coupling constant, α_s , which can be systematically calculated in perturbative QCD.

This paper documents the first release of the open-source code kolya [4]. It consists of a Python library which computes the prediction for the total rate and lepton-energy, hadronic invariant mass and dilepton invariant mass kinematic moments within the framework of the HQE utilizing the so-called kinetic scheme [5–8]. The kolya code supersedes and extends the code developed for the fit of q^2 moments [9] measured by Belle [10] and Belle II [11]. We also include effects from new physics (NP) studied in Ref. [12]. These are parameterised as Wilson coefficients of dimension-six operators in the weak effective theory below the electroweak scale [13, 14].

Several building blocks necessary for the prediction of $\Gamma_{\rm sl}$ and the moments to high orders in α_s and the $1/m_b$ expansion have been presented over the last 30 years (see Sec. 3.2 for an exhaustive list of references). The kolya library provides the first comprehensive open-source framework in which all available corrections are implemented and validated. A schematic overview of the perturbative corrections implemented for the total rate and the moments is given in Tab. 1. This document accompanies the first release of kolya and details the specifics of the code. Although this paper represents a reference for future analyses of Belle II measurements of inclusive $B \to X_c l \bar{\nu}_l$ decays and gives a first outlook of kolya with basic examples to try in a Jupyter notebook, it is not meant to be a review article on semileptonic decays. To obtain a deeper understanding of the scientific part, the user is referred to e.g. Refs. [15–17]. The software kolya complements in scope several other open-source packages in HEP, in particular flavio [18], EOS [19], HEPfit [20], HAMMER [21] and SuperIso [22].

This article is structured as follows. In Sec. 2, we present the definitions of observables. Their implementation in the code is discussed in Sec. 3 where we discuss various ingredients implemented and quote the original references from which the material was obtained. The definition of the effective Hamiltonian parametrizing NP effects is given in Sec. 4. Section 5 focuses on basic usage of the code, illustrating the installation, the classes implemented in the code, use of the code for calculating $\Gamma_{\rm sl}$ and the moments together with details about our validation of the code. We close in Sec. 6 with an outlook.

2 Definitions

We consider the semileptonic decay

$$\overline{B}(p_B) \to X_c(p_X)l(p_l)\bar{\nu}_l(p_\nu) \text{ with } l = e, \mu,$$
 (1)

and describe the decay rate in the rest frame of the B meson, i.e. $p_B = (M_B, \vec{0})$. The leptons are considered massless. We denote the total momentum of the lepton pair by $q = p_l + p_{\nu}$, the total momentum of the hadronic system by $p_X = p_B - q$, and the electron energy by E_l .

Within the HQE, it is possible to make a prediction for the various differential rates w.r.t. the q^2 , the total leptonic energy $q_0 = E_{\nu} + E_l$, and the energy E_l of the charged lepton. However, the predictions for the differential rates cannot be compared point by point with data. This is because, on the one hand, the phase space region allowed at the parton level is smaller than the physical one. On the other hand, power corrections become singular close to the endpoint. Instead, data of inclusive $\overline{B} \to X_c l \bar{\nu}_l$ has to be compared to theory predictions of integrated quantities like the total rate

$$\Gamma_{\rm sl} = \int \frac{\mathrm{d}^3 \Gamma}{\mathrm{d}q^2 \,\mathrm{d}q_0 \,\mathrm{d}E_l} \mathrm{d}q^2 \,\mathrm{d}q_0 \,\mathrm{d}E_l \tag{2}$$

or moments of the differential distribution of some observable O, where $O = E_l, q^2, M_X^2$, normalized w.r.t. the partial decay width. The moments are defined by

$$\langle (O)^n \rangle_{\text{cut}} = \int_{\text{cut}} (O)^n \frac{d\Gamma}{dO} dO / \int_{\text{cut}} \frac{d\Gamma}{dO} dO ,$$
 (3)

where $d\Gamma/dO$ is the differential rate for the variable O. The subscript "cut" generically denotes some restriction in the lower integration limit. From the theoretical side, the dependence of the moments on a lower cut yields additional information on the HQE parameters and thus provides a better handle for their extraction via global fits. From the experimental side, the spectrum is usually not measurable entirely due to detector acceptance. For example at the B-factories, a lower cut on the charged lepton energy, $E_{\ell} \geq E_{\text{cut}}$ with $E_{\text{cut}} \simeq 0.5$ GeV, is applied to suppress the background. It is also possible to consider the partial decay width with a cut on E_{ℓ} or q^2 , defined by

$$\Delta\Gamma_{\rm sl}(E_{\rm cut}) = \int_{E_l > E_{\rm cut}} \frac{\mathrm{d}\Gamma}{\mathrm{d}E_l} \, \mathrm{d}E_l \,, \qquad \Delta\Gamma_{\rm sl}(q_{\rm cut}^2) = \int_{q^2 > q_{\rm cut}^2} \frac{\mathrm{d}\Gamma}{\mathrm{d}q^2} \, \mathrm{d}q^2. \tag{4}$$

For higher moments $(n \ge 2)$, usually centralized moments are considered. The centralized moments of the charged lepton energy are then defined as:

$$\ell_1(E_{\text{cut}}) = \langle E_l \rangle_{E_l \ge E_{\text{cut}}}, \qquad \ell_n(E_{\text{cut}}) = \langle (E_l - \langle E_l \rangle)^n \rangle_{E_l \ge E_{\text{cut}}} \text{ for } n \ge 2, \qquad (5)$$

and moments of the hadronic invariant mass as:

$$h_1(E_{\text{cut}}) = \langle M_X^2 \rangle_{E_l \ge E_{\text{cut}}}, \qquad h_n(E_{\text{cut}}) = \left\langle (M_X^2 - \langle M_X^2 \rangle)^n \right\rangle_{E_l \ge E_{\text{cut}}} \text{ for } n \ge 2.$$
 (6)

The first moments with n = 1 correspond to the mean value of observable over the considered integration domain, while the second centralized moments are the variances of the distributions. Experimentally, moments up to n = 3 and n = 4 are currently available. Ref. [23] proposed also the study of the moments of the q^2 spectrum, defined by

$$q_1(q_{\text{cut}}^2) = \langle q^2 \rangle_{q^2 \ge q_{\text{cut}}^2}$$
, $q_n(q_{\text{cut}}^2) = \left\langle (q^2 - \langle q^2 \rangle)^n \right\rangle_{q^2 \ge q_{\text{cut}}^2}$ for $n \ge 2$. (7)

These q^2 moments are invariant under reparametrization and therefore they depend on a reduced set of HQE parameters (which we will introduce in Sec. 3), like the total rate (see Refs. [23, 24]). Since a cut on E_l in the definition of the q^2 moments would break reparametrization invariance (RPI), Ref. [23] suggested to consider instead a lower cut on q^2 . Such a lower cut on q^2 also imposes an indirect cut on the charged lepton energy through

$$E_l \ge \frac{M_B^2 + q_{\text{cut}}^2 - M_D^2 - \lambda^{1/2}(M_B^2, q_{\text{cut}}^2, M_D^2)}{2M_B} , \qquad (8)$$

where $\lambda(a,b,c) = a^2 + b^2 + c^2 - 2ab - 2ac - 2bc$ is the Källén function. A cut on q^2 is therefore capable of excluding low-energy electrons in the experimental analysis, on equal footing as a cut on E_l .

3 Implementation in the SM

3.1 Building blocks

We use the heavy quark expansion (HQE) and express the total semileptonic width $\Gamma_{\rm sl}$ and the kinematic moments as a double expansion in $1/m_b$ and α_s . While working with the HQE, it is often advantageous to consider dimensionless quantities normalized w.r.t. the bottom quark mass m_b . We will denoted them with a "hat(^)": e.g. $\hat{q}^2 = q^2/m_b^2$, $\hat{E}_l = E_l/m_b$.

As a starting point, we define the following building blocks:

$$Q_{ij} = \frac{1}{\Gamma_0} \int_{\text{cut}} dE_l \, dq_0 \, dq^2 \, (q^2)^i (q_0)^j \, \frac{d^3 \Gamma}{dq^2 \, dq_0 \, dE_l} \,, \tag{9}$$

where $q_0 = (E_l + E_{\nu}) = v \cdot q$ is the total leptonic energy with $v = p_B/M_B$ and q^2 is the

leptonic invariant mass. Schematically, we write

$$Q_{ij} = (m_b)^{2i+j} \left[Q_{i,j}^{(0)} + Q_{i,j}^{(1)} \frac{\alpha_s(\mu_s)}{\pi} + Q_{i,j}^{(2)} \left(\frac{\alpha_s(\mu_s)}{\pi} \right)^2 + \frac{\mu_\pi^2}{m_b^2} \left(Q_{i,j,\pi}^{(0)} + Q_{i,j,\pi}^{(1)} \frac{\alpha_s(\mu_s)}{\pi} \right) \right.$$

$$\left. \frac{\mu_G^2(\mu_b)}{m_b^2} \left(Q_{i,j,G}^{(0)} + Q_{i,j,G}^{(1)} \frac{\alpha_s(\mu_s)}{\pi} \right) + \frac{\rho_D^3(\mu_b)}{m_b^3} \left(Q_{i,j,D}^{(0)} + Q_{i,j,D}^{(1)} \frac{\alpha_s(\mu_s)}{\pi} \right) \right.$$

$$\left. + \frac{\rho_{LS}^3(\mu_b)}{m_b^3} \left(Q_{i,j,LS}^{(0)} + Q_{i,j,LS}^{(1)} \frac{\alpha_s(\mu_s)}{\pi} \right) + O\left(\frac{1}{m_b^4} \right) \right], \tag{10}$$

where

$$\Gamma_0 = \frac{m_b^5 G_F^2 A_{\text{ew}}}{192\pi^3} |V_{cb}|^2 \ . \tag{11}$$

The factor $A_{\rm ew}=1.01435$ stems from short-distance radiative corrections at the electroweak scale [25]. $\alpha_s \equiv \alpha_s^{(n_f)}(\mu_s)$ is the strong coupling constant taken with n_f active quarks and at the renormalization scale μ_s . To leading order in $1/m_b$, the heavy B meson decay coincides with the decay of a free bottom computed in perturbative QCD. Starting from $O(1/m_b^2)$ the predictions depend on a set of HQE parameters: non-perturbative matrix elements of local operators. These are denoted by $\mu_\pi^2, \mu_G^2, \rho_D^3, \rho_{LS}^3$. The tree-level expressions are known also to higher orders in $1/m_b$ (see Refs. [23, 26–28]). They are implemented in kolya up to $1/m_b^5$. However, in (10) they are omitted to keep a compact notation. The explicit definitions of the HQE parameters up to $1/m_b^5$ are reported in Appendix A. The HQE parameters in (10) are quoted in what we refer to as the "historical" basis employed in e.g. [29, 30]. For RPI quantities, like q^2 moments, it is however useful to work in the RPI basis, which has a reduced number of parameters [23, 31]. The differences between these two bases are detailed in Ref. [31, 32].

The functions denoted by Q_{ij} are the fundamental building blocks necessary to assemble the predictions for the centralized moments q_i in (7) and h_i in (6). They all depend on the mass ratio

$$\rho \equiv \frac{m_c}{m_b} \,\,, \tag{12}$$

where m_c and m_b refer to the on-shell masses of the charm and bottom quark. In (9), the subscript "cut" refers to certain restrictions in the phase-space integration. For the prediction of q_i in (7), we apply the cut $q^2 > q_{\text{cut}}^2$ so that various build blocks in (10) depend on ρ and \hat{q}_{cut}^2 : $Q_{ij} \to Q_{ij}(\rho, \hat{q}_{\text{cut}}^2)$. For the hadronic moments h_n in (6), the restriction is on the electron energy $E_l > E_{\text{cut}}$, so that the building blocks are functions of ρ and \hat{E}_{cut} : $Q_{ij} \to Q_{ij}(\rho, \hat{E}_{\text{cut}})$ (see Eq. (18) for the relation between h_n and Q_{ij}).

The QCD corrections depend also on the renormalization scale μ_s of the strong coupling constant starting at $O(\alpha_s^2)$. The functions $Q_{i,j,G}$, $Q_{i,j,D}$ and $Q_{i,j,LS}$ depend on μ_b , the scale at which the Wilson coefficients of the HQET Lagrangian are matched onto QCD, starting from $O(\alpha_s)$.

To construct the centralized electron energy moments in (5), we consider the moments of the charged-lepton energy $E_l = p_l \cdot v$ within the HQE:

$$L_{i} = \frac{1}{\Gamma_{0}} \int_{E_{l} \geq E_{\text{cut}}} dE_{l} dq_{0} dq^{2} (E_{l})^{i} \frac{d^{3}\Gamma}{dq^{2} dq_{0} dE_{l}}$$

$$= (m_{b})^{i} \left[L_{i}^{(0)} + L_{i}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} + L_{i}^{(2)} \left(\frac{\alpha_{s}(\mu_{s})}{\pi} \right)^{2} + \frac{\mu_{\pi}^{2}}{m_{b}^{2}} \left(L_{i,\pi}^{(0)} + L_{i,\pi}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} \right) \right.$$

$$+ \frac{\mu_{G}^{2}(\mu_{b})}{m_{b}^{2}} \left(L_{i,G}^{(0)} + L_{i,G}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} \right) + \frac{\rho_{D}^{3}(\mu_{b})}{m_{b}^{3}} \left(L_{i,D}^{(0)} + L_{i,D}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} \right)$$

$$+ \frac{\rho_{LS}^{3}(\mu_{b})}{m_{b}^{3}} \left(L_{i,LS}^{(0)} + L_{i,LS}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} \right) + O\left(\frac{1}{m_{b}^{4}}\right) \right], \tag{13}$$

where in this case we allow for a cut on E_l only. All functions L_i depend on ρ and the cut E_{cut} : $L_i \to L_i(\rho, \hat{E}_{\text{cut}})$.

The total semileptonic rate corresponds to

$$\Gamma_{\rm sl} = \frac{m_b^5 G_F^2 A_{\rm ew}}{192\pi^3} |V_{cb}|^2 Q_{0,0}(\rho,0) = \Gamma_0 L_0(\rho,0), \tag{14}$$

with no cut applied, namely $E_{\rm cut}=q_{\rm cut}^2=0$. For the partial decay width, we similarly have

$$\Delta\Gamma_{\rm sl}(E_{\rm cut}) = \Gamma_0 L_0(\rho, \hat{E}_{\rm cut}) , \qquad \Delta\Gamma_{\rm sl}(q_{\rm cut}^2) = \Gamma_0 Q_{0.0}(\rho, \hat{q}_{\rm cut}^2) . \qquad (15)$$

The ratios defined in (3) correspond to

$$\langle (q^2)^n \rangle = \frac{Q_{n,0}}{Q_{0,0}} , \qquad \langle E_l^n \rangle = \frac{L_n}{L_0} . \qquad (16)$$

The centralized moments are obtained by inserting the double expansions of (10) or (13) into (3-7) and re-expanding in α_s and $1/m_b$ up to the relevant order. To assemble the M_X^2 moments, we express the hadronic invariant mass in terms of the parton level quantities in the B rest frame:

$$M_X^2 = (M_B v - q)^2 = M_B^2 + q^2 - 2M_B q_0. (17)$$

The moments of M_X^2 are obtained as linear combinations of the mixed moments $Q_{i,j}$:

$$M_{n} = \frac{1}{\Gamma_{0}} \int_{E_{l} \geq E_{\text{cut}}} dE_{l} dq_{0} dq^{2} (M_{B}^{2} - 2M_{B}q_{0} + q^{2})^{n} \frac{d^{3}\Gamma}{dE_{l} dq_{0} dq^{2}}$$

$$= \sum_{i=0}^{n} \sum_{j=0}^{i} \binom{n}{i} \binom{i}{j} (M_{B}^{2})^{n-i} (-2M_{B})^{i-j} Q_{j,i-j} , \qquad (18)$$

$\Gamma_{ m sl}$	tree	α_s	α_s^2	α_s^3
Partonic		[34]	[35–38]	[39]
μ_π^2, μ_G^2	[1, 2]	[40-43]		
$ ho_D^3, ho_{LS}^3$	[44]	[45]		
$1/m_b^4, 1/m_b^5$	[23, 26–28]			
$q_n(q_{\mathrm{cut}}^2)$	tree	α_s	α_s^2	
Partonic		[45, 46]	[47]	
μ_G^2, μ_π^2	[1, 2]	[41, 42]		
$ ho_D^3, ho_{LS}$	[44]	[45]		
$1/m_b^4, 1/m_b^5$	[23, 28]			
$\ell_n(E_{\mathrm{cut}}), h_n(E_{\mathrm{cut}})$	tree	α_s	$\alpha_s^2 \beta_0$	α_s^2
Partonic		[46, 48, 49]	[46]	[33]*
μ_G^2, μ_π^2	[1, 2]	[42, 50]		
$ ho_D^3$	[44]			
$1/m_b^4, 1/m_b^5$	[26-28]			

Table 1: Schematic overview of the perturbative corrections implemented for the rate $\Gamma_{\rm sl}$, the q^2 moments, the E_l and M_X^2 moments. (*) The α_s^2 corrections to h_n and ℓ_n are only available for several ρ and $E_{\rm cut}$ values in [33].

and
$$\langle (M_X^2)^n \rangle = M_n/M_0$$
.

In kolya, we first implement all building blocks introduced above, corresponding to the on-shell scheme for both m_b and m_c . We collect in the Tab. 1 the list of references from which the various building blocks are retrieved. The implementation of the building blocks is described in Sec. 3.2. The implementation of the NNLO corrections to E_l and M_X moments, based on the results published in Ref. [33], requires a dedicated discussion in sections 3.3 and 3.4. In Sec. 3.5, we perform a scheme change to the kinetic scheme to obtain the final prediction for the total rate and the centralized moments.

3.2 Analytic expressions and grids for QCD corrections

In kolya, the tree level expressions up to $O(1/m_b^5)$ (see Refs. [23, 26–28]) are implemented in an exact analytic form. For example, the tree-level expressions at leading order in $1/m_b$ for $L_i^{(0)}(\rho, \hat{E}_{\text{cut}})$ in (13) are coded in Python as follows.

from numba import jit
import math

```
@jit(cache=True, nopython=True)
def L_0(i,elcuthat,r,dEl,dr):
    """ Tree level (partonic) for El moments and their derivatives """
    y = 2*elcuthat
    logy = math.log((1-y)/r**2)
    # tree level function
    if (dEl == 0 \text{ and } dr == 0 \text{ and } i==0):
        return (1-8*r**2-6*r**4+12*logy*r**4+4*r**6-r**8
             -(2*r**6)/(-1+y)**2-(6*r**4*(1
            +r**2))/(-1+y)
             -2*r**4*(-3+r**2)*y+2*(-1+r**2)*y**3+y**4)
    if (dEl == 0 \text{ and } dr == 0 \text{ and } i==1):
        return (3*logy*r**4*(3+r**2)+(7-75*r**2-180*r**4
             +120*r**6-15*r**8+3*r**10)/20
             -r**6/(-1+y)**2-(r**4*(3+5*r**2))/(-1+y)
             +6*r**4*y-(r**4*(-3+r**2)*y**2)/2
             +(3*(-1+r**2)*y**4)/4+(2*y**5)/5)
```

where the arguments of L_0 refer to the moment $i \in [0, ..., 4]$, the normalized electron energy cut $\hat{E}_{\rm cut} = E_{\rm cut}/m_b$ (elcuthat) and the mass ratio $\rho = m_c/m_b$ indicated by (r). The additional two arguments (dEl,dr) are positive integers referring to the derivatives of $L_i(\rho, \hat{E}_{\rm cut})$ w.r.t. to $\hat{E}_{\rm cut}$ or ρ . These derivatives are required when expressing the predictions in the kinetic scheme (see detailed discussion in Sec. 3.5). The tree level expressions up to $O(1/m_b^3)$ for the moments are implemented in the files Q2moments_SM.py, Elmoments_SM.py and MXmoments_SM.py. The power corrections of order $1/m_b^4$ and $1/m_b^5$ are given in separate files Q2moments_HO.py, Elmoments_HO.py and MXmoments_HO.py. For the q^2 moments, the higher power corrections are in Q2moments_HO_RPI.py

In order to use kolya for phenomenological applications, for instance to perform a fit, it is important to ensure adequate speed for the numerical evaluations. To this end, our implementation utilizes the library Numba [51], which translates Python functions to optimized machine code at runtime using the standard LLVM compiler library. The functions decorated with @jit, as shown in the example above, are compiled to machine code "just-in-time" for execution at native machine-code speed. Numba-compiled routines in Python approach the speed of C or FORTRAN.

Let us now discuss the implementation of the QCD corrections in kolya. For $\Gamma_{\rm sl}$, the functions in (10) depend only on the mass ratio ρ and we implement the NLO corrections up to $1/m_b^3$ using the analytic expressions given in Ref. [45]. At NNLO in the free bottom quark approximation, there are asymptotic expansions available either in the limit $\rho \to 0$ [35, 36] or $\delta = 1 - \rho \to 0$ [37]. Recently, analytic expressions for the NNLO corrections written in terms of iterated integrals were presented in Ref. [38]. In kolya we use the expressions expanded in terms of $\delta = 1 - \rho$ up to δ^{46} given in Ref. [38]. For the third-order correction to $\Gamma_{\rm sl}$ we implement the asymptotic expansion up to δ^{12} computed in Ref. [39]. The total rate is implemented up to $1/m_b^3$ in TotalRate_SM.py, the higher power corrections are given in

TotalRate_HO.py and TotalRate_HO_RPI.py.

We use interpolation grids to implements most of the QCD corrections for the moments. Specifically, we use grids for all NLO corrections, the NNLO corrections to the q^2 moments and the so-called BLM corrections (of order $\alpha_s^2\beta_0$) [52] to the E_l and M_X moments.

Analytic expressions for the differential rate $d\Gamma/dq^2$ for a free bottom quark are available at NLO from Ref. [53] and at NNLO from Ref. [47]. The NLO corrections to leading order in $1/m_b$ were computed also for the triple differential rate in Ref. [46, 48]. The NLO corrections to μ_G^2 and ρ_D^3 for the q^2 spectrum have been presented in Ref. [45]. By making use of reparametrization invariance [24], one can also show that in the on-shell scheme

$$-\frac{1}{2}Q_{i0}^{(n)}(\rho, q_{\text{cut}}^2) = Q_{i0,\pi}^{(n)}(\rho, q_{\text{cut}}^2) , \qquad -Q_{i0,G}^{(n)}(\rho, q_{\text{cut}}^2) = Q_{i0,LS}^{(n)}(\rho, q_{\text{cut}}^2) , \qquad (19)$$

to all orders $n \geq 0$ in the perturbative expansion. Therefore, the functions Q_{ij} entering (10) can be computed for any q_{cut}^2 and ρ via:

$$Q_{i0}(\rho, \hat{q}_{\text{cut}}^2) = \frac{1}{\Gamma_0} \int_{\hat{q}_{\text{cut}}^2}^{(1-\rho)^2} (\hat{q}^2)^i \frac{d\Gamma}{d\hat{q}^2} d\hat{q}^2 .$$
 (20)

The differential rate $d\Gamma/d\hat{q}^2$ at higher orders is expressed in terms of functions defined via iterated integrals like harmonic polylogarithms (HPLs) [54] and generalized polylogarithms (GPLs) [55, 56]. It is not convenient to integrate the differential rate numerically "on-the-fly" since there are several HPLs and GPLs whose evaluation (for instance with GiNaC [57]) is time-consuming. For this reason, we opt to implement all higher QCD corrections for the moments not in an exact form, but through Chebyshev two-dimensional grids. The functions implemented via grids are $Q_{ij}^{(1)}, Q_{ij}^{(2)}, Q_{ij,\pi}^{(1)}Q_{ij,G}^{(1)}, Q_{ij,D}^{(1)}, Q_{ij,LS}^{(1)}$ in (10) and $L_{ij}^{(1)}, L_{ij}^{(2)}, L_{ij,\pi}^{(1)}L_{ij,G}^{(1)}$ in (13). Let us briefly review here how a generic function f(x) can be discretized on a grid

Let us briefly review here how a generic function f(x) can be discretized on a grid consisting of the so-called Chebyshev points (for more details see e.g. Ref. [58]). The idea is to evaluate f(x) in n points x_n corresponding to the zeros of the Chebyshev polynomial $T_n(x)$ of degree n:

$$x_k = \cos\left(\frac{\pi(k-1/2)}{n}\right) \text{ for } k = 1,\dots, n.$$
 (21)

If f(x) is an arbitrary function defined on the domain $x \in [-1, 1]$, we calculate the coefficients c_j , with j = 0, ..., n - 1 given by

$$c_j = \frac{2}{n} \sum_{k=1}^n f(x_k) T_j(x_k) , \qquad (22)$$

which can be employed to construct the polynomial

$$\tilde{f}(x) = -\frac{1}{2}c_0 + \sum_{k=1}^{n-1} c_k T_k(x) \approx f(x) , \qquad (23)$$

approximating f(x) in the interval [-1,1]. In particular $\tilde{f}(x) = f(x)$ for all n zeros of $T_n(x)$. In case the function to interpolate f(y) is defined between two arbitrary limits, e.g. $y \in [a,b]$, we apply the variable transformation

$$x = \frac{y - \frac{1}{2}(b+a)}{\frac{1}{2}(b-a)} , \qquad (24)$$

and then perform the interpolation in x as before.

In our setup, the functions to interpolate depend on ρ and \hat{q}_{cut}^2 (or ρ and \hat{E}_{cut}) and can be obtained via two consecutive one-dimensional Chebyshev interpolations. First, we discretize the interval $\rho \in [1/6, 1/3]$ (relevant for the phenomenology) in n_{ρ} points distributed according to (21). Then for each $\rho_k \in \{\rho_1, \dots, \rho_{n_{\rho}}\}$, we discretize \hat{q}_{cut}^2 or \hat{E}_{cut} into further n_{cut} points within the allowed range: $\hat{q}_{\text{cut}}^2 \in [0, (1-\rho)^2]$ or $\hat{E}_{\text{cut}} \in [0, (1-\rho^2)/2]$. An example of how the discretization is performed is shown in Fig. 1, for a grid in ρ and \hat{q}_{cut}^2 with $n_{\rho} = n_{\text{cut}} = 10$.

To estimate the function at a new point $P = (\rho_P, \hat{q}_P^2)$ (the green diamond in Fig. 1), we proceed as follows. For each $\rho_k, k = 1, ..., n_\rho$, we calculate $Q(\rho_k, \hat{q}_P^2(1 - \rho_P)^2/(1 - \rho_k)^2)$ using one-dimensional interpolations in the variable \hat{q}_{cut}^2 . These values calculated at fixed ρ are shown by black crosses in Fig. 1. Afterwards, we use them as nodes for a second interpolation this time along the ρ direction, as displayed by the black dashed line in Fig. 1. The second interpolation yields the final estimate for $f(\rho_P, \hat{q}_P^2)$.

For the implementation of the QCD corrections to h_i and ℓ_i , which depend on ρ and $E_{\rm cut}$, we also use Chebyshev interpolation grids. At NLO, it is possible to write the differential rate $d\Gamma/dE_l$ in a closed analytic form at NLO [49] for a free quark. To compute $Q_{ij}(\rho, \hat{E}_{\rm cut})$ and $L_i(\rho, \hat{E}_{\rm cut})$ at $O(\alpha_s)$ we perform the one variable integration, as for instance

$$L_{i}(\rho, \hat{E}_{\text{cut}}) = \frac{1}{\Gamma_{0}} \int_{\hat{E}_{\text{cut}}}^{(1-\rho^{2})/2} (\hat{E}_{l})^{i} \frac{d\Gamma}{d\hat{E}_{l}} d\hat{E}_{l} ,$$

$$Q_{ij}(\rho, \hat{E}_{\text{cut}}) = \frac{1}{\Gamma_{0}} \int_{\hat{E}_{\text{cut}}}^{(1-\rho^{2})/2} \frac{dQ_{ij}}{d\hat{E}_{l}} d\hat{E}_{l} .$$

$$(25)$$

In the last equation, we define

$$\frac{\mathrm{d}Q_{ij}}{\mathrm{d}\hat{E}_l} = \int (q^2)^i (q_0)^j \frac{\mathrm{d}^3 \Gamma}{\mathrm{d}q^2 \,\mathrm{d}q_0 \,\mathrm{d}El} \,\mathrm{d}q^2 \mathrm{d}q_0 \tag{26}$$

which can also be computed analytically up to NLO following Ref. [49, 59].

At order $1/m_b^2$, the NLO corrections $Q_{ij}(\rho, \bar{E}_{\text{cut}})$ and $L_i(\rho, \bar{E}_{\text{cut}})$ can be calculated from the triple differential distributions given in Refs. [41, 42] by performing the phase space integration numerically as described in Ref. [60]. The NLO corrections to ρ_D^3 and ρ_{LS}^3 are not known at the moment.

The values of the coefficients c_j in (22) for all grids are stored in the directory grids as multidimensional arrays. The routines which perform the interpolation of the NLO and

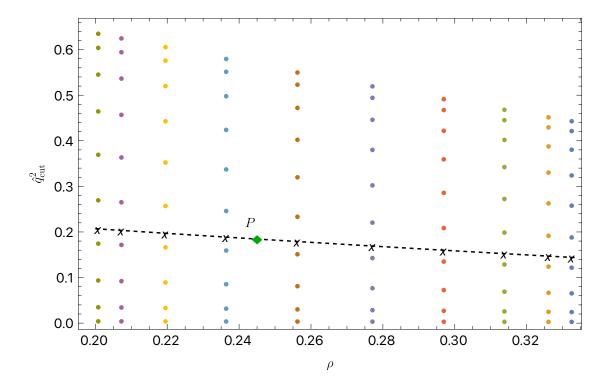


Figure 1: Example of how we discretize functions depending on two variables using Chebyshev nodes. In the plot, we consider the grid for Q_{ij} depending on ρ and \hat{q}_{cut}^2 .

NNLO corrections at $O(1/m_b^0)$ are implemented in NLOpartonic.py and NNLOpartonic.py. The routines for the NLO corrections to the power-suppressed terms are given in NLOpw.py.

We validate the grid implementation by generating 100 random points in the two-dimensional plane $(\rho, \hat{q}_{\text{cut}}^2)$ or $(\rho, \hat{E}_{\text{cut}})$. For each point, we compare the approximation provided by the grids and high-precision evaluations obtained with Mathematica. We verify that the two estimates differ by less than 10^{-5} for the points considered.

3.3 NNLO corrections to the lepton energy moments

The NNLO corrections to the ℓ_i moments are not known in a closed form. As discussed, the BLM corrections are implemented through interpolation grids. The remaining "non-BLM" corrections are only known for specific values of ρ and $E_{\rm cut}$ from Ref. [33]. Their functional form can be obtained from a two-dimensional fit to these data points. In order to perform this fit, we write:

$$\ell_{n}(\rho, E_{\text{cut}}) = (m_{b})^{n} \left[Y_{n}^{(0)} + Y_{n}^{(1)} \frac{\alpha_{s}(\mu_{s})}{\pi} + \left(\beta_{0} Y_{n}^{(2,\text{BLM})} + Y_{n}^{(2,\text{nonBLM})} \right) \left(\frac{\alpha_{s}(\mu_{s})}{\pi} \right)^{2} + O\left(\frac{1}{m_{b}}, \left(\frac{\alpha_{s}(\mu_{s})}{\pi} \right)^{3} \right) \right],$$
(27)

where the E_{cut} - and ρ -dependence of $Y_n^{(i)}$ is implied and $Y_n^{(0)}$ is the partonic contribution without any α_s or $1/m_b$ corrections. In terms of the building-blocks defined in (13), we can write the non-BLM terms as:

$$Y_{1}^{(2,\text{nonBLM})} = \frac{L_{1}^{(2,\text{nonBLM})}}{L_{0}^{(0)}} - \frac{L_{0}^{(2,\text{nonBLM})}L_{1}^{(0)}}{(L_{0}^{(0)})^{2}},$$

$$Y_{2}^{(2,\text{nonBLM})} = \frac{L_{2}^{(2,\text{nonBLM})}}{L_{0}^{(0)}} - 2\frac{L_{1}^{(2,\text{nonBLM})}L_{1}^{(0)}}{(L_{0}^{(0)})^{2}} + L_{0}^{(2,\text{nonBLM})}\left(2\frac{(L_{1}^{(0)})^{2}}{(L_{0}^{(0)})^{3}} - \frac{L_{2}^{(0)}}{(L_{0}^{(0)})^{2}}\right),$$

$$Y_{3}^{(2,\text{nonBLM})} = \frac{L_{3}^{(2,\text{nonBLM})}}{L_{0}^{(0)}} - 3\frac{L_{2}^{(2,\text{nonBLM})}L_{1}^{(0)}}{(L_{0}^{(0)})^{2}} + 3L_{1}^{(2,\text{nonBLM})}\left(2\frac{(L_{1}^{(0)})^{2}}{(L_{0}^{(0)})^{3}} - \frac{L_{2}^{(0)}}{(L_{0}^{(0)})^{2}}\right)$$

$$+ L_{0}^{(2,\text{nonBLM})}\left(-6\frac{(L_{1}^{(0)})^{3}}{(L_{0}^{(0)})^{4}} + 6\frac{L_{1}^{(0)}L_{2}^{(0)}}{(L_{0}^{(0)})^{2}} - \frac{L_{3}^{(0)}}{(L_{0}^{(0)})^{2}}\right). \tag{28}$$

Ref. [33] gives the $L_n^{(2,\text{nonBLM})}$ terms at $\rho = \{0.20, 0.22, 0.24, 0.25, 0.26, 0.28\}$ and $y \equiv 2\hat{E}_{\text{cut}} = \{0, 0.1, \dots, 0.7\}$. From these¹, the non-BLM contributions $Y_n^{(2,\text{nonBLM})}$ to the ℓ_n moments are obtained by combining with the tree-level building blocks $L_n^{(0)}$. In Ref. [61], these non-BLM contributions are studied in detail and compared to the effect of their BLM counterparts. We fit the values for $Y_n^{(2,\text{nonBLM})}$ assuming the following polynomial ansatz

$$Y_n(\rho, y) = \sum_{i=1}^{5} (a_{n,i} + b_{n,i}\rho)(y + \rho^2 - 1)^i , \qquad (29)$$

for each moment n. Following Ref. [61], we only include one power of ρ in our ansatz, but keep up terms up y^5 in our interpolating fit. In addition, the ansatz is chosen to ensure that the non-BLM corrections vanish at the end point $y = 1 - \rho^2$ [33]. We stress that our approach differs from [61] as we first construct $Y_n^{(2,\text{nonBLM})}$ in each available (ρ, y) point and then perform the analysis. Fitting first the $L_n^{(2,\text{nonBLM})}$ and then combining them resulted

Note that the $L_i^{(n)}$ defined in Ref. [33] are normalized to the total partonic rate without cut while we only normalize to Γ_0 defined in (14).

in strongly oscillating functions due to accidental cancellations. Fitting directly Y_n , we find

$$Y_{1}^{(2,\text{nonBLM})}(\rho, y) = (72.57\rho - 25.62) (y + \rho^{2} - 1)^{5} + (177.60\rho - 64.65) (y + \rho^{2} - 1)^{4} + (157.19\rho - 59.27) (y + \rho^{2} - 1)^{3} + (62.69\rho - 24.32) (y + \rho^{2} - 1)^{2} + (11.25\rho - 4.35) (y + \rho^{2} - 1) ,$$

$$Y_{2}^{(2,\text{nonBLM})}(\rho, y) = (12.61 - 44.47\rho) (y + \rho^{2} - 1)^{5} + (32.28 - 112.34\rho) (y + \rho^{2} - 1)^{4} + (29.36 - 100.36\rho) (y + \rho^{2} - 1)^{3} + (11.07 - 36.89\rho) (y + \rho^{2} - 1)^{2} + (1.42 - 4.55\rho) (y + \rho^{2} - 1) ,$$

$$(30)$$

and

$$Y_3^{(2,\text{nonBLM})}(\rho, y) =$$

$$(32.23\rho - 7.28) (y + \rho^2 - 1)^5 + (79.34\rho - 17.72) (y + \rho^2 - 1)^4$$

$$+ (67.06\rho - 14.67) (y + \rho^2 - 1)^3 + (21.52\rho - 4.49) (y + \rho^2 - 1)^2$$

$$+ (1.71\rho - 0.289) (y + \rho^2 - 1) .$$

$$(31)$$

These functions are implemented in kolya. We do not assign any uncertainty to these fit functions.

In Fig. 2, we show our fit results for $Y_{1,2,3}^{(2,\text{nonBLM})}$, as a function of y for fixed $\rho = 0.22$ (solid blue line). Due to the fit ansatz, we observe a light oscillatory behavior as a function of y. In black, we show the constructed data points at fixed y obtained from [33]. The fit uncertainty is given by the red dotted line, which represents the 90% C.L. interval of the fit. Since we only have data points up to y = 0.7 and impose that the contribution vanishes at the endpoint, we notice large uncertainties towards higher y values. In a typical phenomenological analysis, missing higher order terms (here α_s^3) terms would be accounted for by varying the scale of α_s . The blue bland corresponds to the effect of such a scale variation from $[\alpha_s(m_b^{\rm kin}/2), \alpha_s(2m_b^{\rm kin})]$. For Y_1 , we observe that the α_s variation covers the data points and fit uncertainty. For the higher moments, we observe that the fit uncertainty is higher than the α_s variation for large y. However, given the smallness of the BLM contributions to these moments, our default setting is to not include an additional uncertainty for these corrections.

3.4 NNLO corrections to the hadronic invariant mass moments

The nonBLM corrections to the hadronic moments can be obtained from Ref. [46]. However, as for the lepton energy moments, the additional "non-BLM" contributions are only

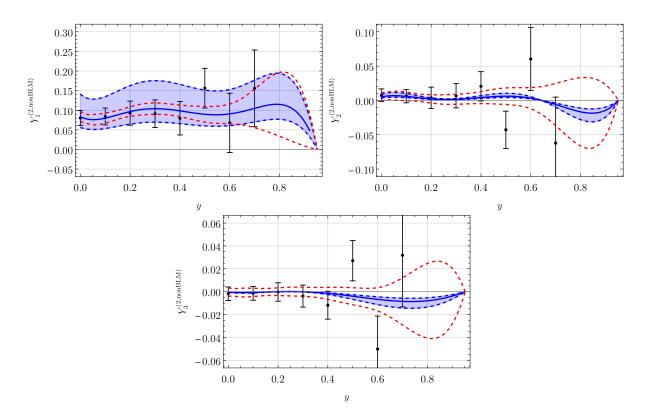


Figure 2: Results of our fitted non-BLM contributions $Y_n^{(2,\text{nonBLM})}$ (solid blue line) as a function of $y = 2E_{\text{cut}}/m_b$ for $\rho = 0.22$ as discussed in the text. Data points are constructed from [33].

known numerically for several values of ρ and $E_{\rm cut}$ [33]. In Ref. [33], building blocks H_{ij} are defined as

$$H_{ij}(\rho, \hat{E}_{\text{cut}}) = \frac{1}{\Gamma_0 L_0^{(0)}(\rho, 0)} \int_{E_{\text{cut}}} d\hat{E}_l \, d\hat{q}_0 \, d\hat{q}^2 \, \left(\frac{m_x^2 - m_c^2}{m_b^2}\right)^i \left(\frac{E_h}{m_b}\right)^j \frac{d^3 \Gamma}{d\hat{E}_l \, d\hat{q}_0 \, d\hat{q}^2} , \qquad (32)$$

where $L_0^{(0)}(\rho,0)$ is defined in (13) and where m_x and E_x are partonic invariant mass and energy. A linear combination of these H_{ij} corresponds to our building blocks Q_{ij} defined in (18). This then allows us to calculate the non-BLM contributions to h_1 by combining different H_{ij} contributions. However, not all the Q_{ij} moments necessary to construct the hadronic moments h_2 and h_3 are calculated. In terms of the definition of [33], the non-BLM contributions of the combinations

$$H_{20}, H_{30}, H_{21}$$

are missing.

In order to determine the effect of the non-BLM terms, we write

$$h_n(y,\rho) = X_n(y,\rho) + m_b^2 \left(\frac{\alpha_s(\mu_s)}{\pi}\right)^2 \left[\beta_0 X_n^{(2,\text{BLM})}(y,\rho) + X_n^{(2,\text{nonBLM})}(y,\rho)\right] ,$$
 (33)

where $X_n^{(2,\text{nonBLM})}$ is the non-BLM contribution, $X_n^{(2,\text{BLM})}$ the BLM contribution and X_n contains all other contributions. For h_1 , we can now proceed as for the lepton energy moments. Here, we use

$$X_1^{(2,\text{nonBLM})}(\rho, y) = \sum_{i=0}^{5} (a_{1,i} + b_{1,i}\rho)y^i , \qquad (34)$$

as our fit ansatz, where we do not require that the first moment vanishes at the end point. Using the data points from Ref. [33] (to be conservative, we assume a 1% uncertainty on all data points for which no uncertainty was given), we then find

$$X_1^{(2,\text{nonBLM})}(y,\rho) = y^5(33.13 - 97.49\rho) + y^4(224.82\rho - 71.16) + y^3(49.01 - 160.90\rho) + y^2(37.72\rho - 10.92) + y(0.73 - 2.56\rho) + 2.08\rho - 0.92.$$
(35)

In Fig. 3, we show $X_1^{(2,\text{nonBLM})}$ as a function of y. As for the lepton energy moments, the red dotted line shows the fit uncertainty, which now diverges close to the end point. However, as usually the experimental data does not have lepton energy cuts y > 0.8 this feature does not pose a great problem. In addition, we notice the small uncertainty on the data points, clearly much smaller than the effect of the α_s variation (shown by the blue band). As such, we do not implement an additional fit uncertainty on (35) and implement this function into kolya. A similar fit was done in [62], where also ρ^2 terms were taken into account. We

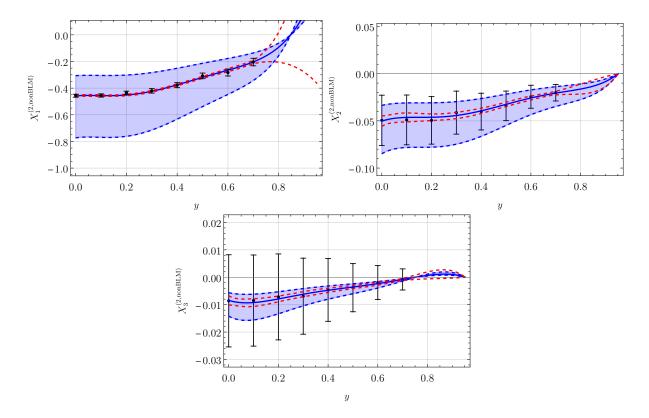


Figure 3: The non-BLM contributions $X_n^{(2,\text{nonBLM})}$ as a function of $y=2E_{\text{cut}}/m_b$ for $\rho=0.22$ as described in the text.

do not notice an improvement in the fit when including such terms and therefore use our minimal fit ansatz presented above.

For the h_2 and h_3 moments, we currently do not have sufficient information to perform a fit. However, for y=0, we can use the expression for Q_{ij} from [59], where analytic expression in terms of $\delta=1-\rho$ were given. From these, we can then also determine the non-BLM effect over the BLM contribution. We find stable results for the different X_n . Using $\rho=0.25$ for illustration, we have

$$\frac{X_1^{(2,\text{nonBLM})}}{\beta_0 X_1^{(2,\text{BLM})}} = -0.237 , \quad \frac{X_2^{(2,\text{nonBLM})}}{\beta_0 X_2^{(2,\text{BLM})}} = -0.227 , \quad \frac{X_3^{(2,\text{nonBLM})}}{\beta_0 X_3^{(2,\text{BLM})}} = -0.241 . \tag{36}$$

We do not quote an uncertainty on these values, which could be obtained from [59] by consider the effect of the highest power in the δ expansion. At the moment, we do not have more information for the X_2 and X_3 contributions at nonzero values of the lepton cut y. In [61], the non-BLM/BLM ratios of the available relevant combinations of H_{ij} moments were calculated. For the values in [33], it was found that this ratio is rather independent of y. This may indicate that also the missing terms are constant when considering ratios of non-BLM versus BLM corrections. In [61], this constant behavior was therefore assumed

for the missing $H_{20,21,30}^{(2)}$ moments². On the other hand, we note that the analytic results in [47] for the q^2 moments with a q^2 cut do not show a constant non-BLM over BLM behavior.

Nevertheless, to further study h_2 and h_3 , we follow [61] and assume constant non-BLM over BLM ratios for the missing moments. We then find the constructed data points for $X_2^{(2,\text{nonBLM})}$ and $X_3^{(2,\text{nonBLM})}$ given in Fig. 3. In blue, we present the fit to the data using a similar ansatz as in (29) (ensuring also that the contribution vanishes at the end point). We note that the fit uncertainty (dotted red line) is very small and that also the α_s variation (blue band) is much smaller than the uncertainty on the data points. In addition, the assumption for the missing moments may bias our predictions. Therefore, we decide for the moment not to include any non-BLM corrections to the h_2 and h_3 hadronic mass moments.

3.5 Kinetic and $\overline{\mathrm{MS}}$ scheme

In the HQE, the expressions for the observables are expressed in terms of a short-distance mass scheme for the bottom and charm quark. This ensures the cancellation of the leading renormalon divergence which arises if the observables are written in terms of the pole mass for the heavy quark [63, 64]. In kolya, we use the so-called kinetic scheme [5] for the bottom quark and the HQE operators. In this scheme, the bottom quark pole mass is rewritten in terms of the kinetic mass using the relation

$$m_b^{\text{pole}} = m_b^{\text{kin}}(\mu) + [\overline{\Lambda}(\mu)]_{\text{pert}} + \frac{[\mu_\pi^2(\mu)]_{\text{pert}}}{2m_b^{\text{kin}}(\mu)} + O\left(\frac{1}{m_b^2}\right), \tag{37}$$

where the scale $\Lambda_{\rm QCD} \ll \mu \ll m_b$ is a Wilsonian cutoff scale, usually taken to be 1 GeV. The latter two terms in (37) denoted with "pert" are the perturbative versions of the HQE parameters determined from the Small Velocity sum rules [65]. Their explicit expressions up to order α_s^3 are given in Refs. [6, 8]. We stress that in our implementation, we follow [8] and include also the decoupling effects of the charm quark in the kinetic scheme. We show the effect of this decoupling numerically in Sec. 3.6. At the same time, the HQE parameters entering the expansion in (10) and (13) must also be redefined in the kinetic scheme by subtracting from their perturbative contribution:

$$\mu_{\pi}^{2}(0) = \mu_{\pi}^{2}(\mu) - [\mu_{\pi}^{2}(\mu)]_{\text{pert}}, \qquad \rho_{D}^{3}(0) = \rho_{D}^{3}(\mu) - [\rho_{D}^{3}(\mu)]_{\text{pert}}.$$
 (38)

These expressions actually refer to the value of μ_{π}^2 and ρ_D^3 in the infinite m_b limit while the fits [9, 29] employ definitions of the HQE parameters at finite m_b . In both cases, the setup neglects (unknown) terms of $O(\mu^3)$ in $[\mu_{\pi}^2(\mu)]_{\rm pert}$ and $[\mu_G^2(\mu)]_{\rm pert}$. Since the operator bases in Refs. [29] and [9] differ in particular for the definition of μ_G^2 , a mismatch of order $\alpha_s \times \mu^3$ appears when comparing the two frameworks. Both bases are implemented in kolya.

Finally, for the charm mass, we use the $\overline{\rm MS}$ scheme. This is related to the pole charm mass via

Using [59], we can calculate for the first time also the ratio of $H_{20}^{(2,\text{non-BLM})}/H_{20}^{(2,BLM)}$ and similarly for $\{21\}$ and $\{30\}$. We confirm within uncertainties the ratios quoted in [61].

$$m_c^{\text{pole}} = \overline{m}_c(\mu_c) \left[1 + \frac{\alpha_s(\mu_c)}{\pi} \left(\frac{4}{3} + L \right) + \left(\frac{\alpha_s(\mu_c)}{\pi} \right)^2 \left(\frac{37}{24} L^2 + \frac{415}{72} L + \frac{779}{96} + \frac{\pi^2}{6} - \frac{\zeta_3}{6} + \frac{1}{9} \pi^2 \log(2) \right) \right] + O(\alpha_s^3), \quad (39)$$

where $L = \log(\mu_c^2/\overline{m}_c^2(\mu_c))$.

In order to implement the scheme change for the quark masses and the HQE parameters, we replace m_b^{pole} and m_c^{pole} in the expressions for the centralized moments using (37) and (39) and expand the formulas as a series in α_s and $1/m_b$.

Hard coding these scheme changes would lead to huge expressions for the moments. To illustrate this, we write the mass conversion formulas schematically as follows:

$$m_c^{\text{pole}} = \overline{m}_c (1 + \alpha_s \delta_{m_c,1} + \alpha_s^2 \delta_{m_c,2} + \alpha_s^3 \delta_{m_c,3}) ,$$

$$m_b^{\text{pole}} = m_b^{\text{kin}} (1 + \alpha_s \delta_{m_b,1} + \alpha_s^2 \delta_{m_b,2} + \alpha_s^3 \delta_{m_b,3}) .$$

$$(40)$$

The coefficients appearing at order α_s^n from the bottom and charm mass scheme conversion formulas are denoted by $\delta_{m_b,n}$ and $\delta_{m_c,n}$, respectively. Let us now consider a simple example with a function $f(\rho)$ depending only on the mass ratio. The scheme change for $f(\rho)$ up to $O(\alpha_s)$ reads

$$f(\rho) = f\left(\frac{m_c^{\text{pole}}}{m_b^{\text{pole}}}\right) = f\left(\overline{m}_c(1 + \alpha_s \delta_{m_c,1}) \middle/ m_b^{\text{kin}}(1 + \alpha_s \delta_{m_b,1})\right)$$
$$= f\left(\frac{\overline{m}_c}{m_b^{\text{kin}}}\right) + \alpha_s f'\left(\frac{\overline{m}_c}{m_b^{\text{kin}}}\right) \frac{\overline{m}_c}{m_b^{\text{kin}}}(\delta_{m_c,1} - \delta_{m_b,1}) + O(\alpha_s^2) . \tag{41}$$

The function $f(\rho)$ and its derivative are the building blocks entering in kolya. This has several advantages with respect to hard coding the expression in.

In fact for the centralized moments, the role of $f(\rho)$ is played by the building blocks denoted by Q_{ij} and L_i on the r.h.s. of (10) and (13). They appear multiple times after the re-expansion of the ratios in (3) and the scheme change. Therefore, we find more convenient to calculate first all building blocks in (10) and (13) (and the necessary derivatives) and cache the results at a given value of ρ and the cut. Afterwards, we assemble the centralized moments using expressions similar to (41). This approach yields code with a much smaller size and improved evaluation time.

The coefficients $\delta_{m_b,n}$ and $\delta_{m_c,n}$ which enter in the scheme change are implemented up to $O(\alpha_s^3)$ in the file schemechange_KINMS.py. This allows in principle to easily adopt a different mass scheme from our default one, by changing the expression for $\delta_{m_b,n}$ and $\delta_{m_c,n}$ to the required mass scheme in a separate file.

Moment	tree	α_s	$\alpha_s^2 \beta_0$	α_s^2 scheme	α_s^2 non-BLM	$\alpha_s^2 \Delta_c$
ℓ_1 [GeV]	1.5650	1.5521	1.5540	1.5459	1.5480	1.5465
$\ell_2 \; [\mathrm{GeV^2}]$	0.0895	0.0870	0.0881	0.0861	0.0867	0.0863
$\ell_3 \; [\mathrm{GeV^3}]$	-0.0018	-0.0003	0.0004	0.0006	-0.0006	-0.0006
$h_1 [\mathrm{GeV^2}]$	4.166	4.331	4.304	4.417	4.381	4.403
$h_2 \; [\mathrm{GeV^4}]$	0.609	0.818	1.001	0.987	-	0.990
$h_3 [\mathrm{GeV^6}]$	5.071	4.810	4.487	4.641	_	4.640

Table 2: Contributions of the different orders to the moments at $E_{\rm cut}=1~{\rm GeV}$ for $\mu_\pi^2=0.4~{\rm GeV}^2$, $\mu_G^2=0.35~{\rm GeV}^2$, $\rho_{LS}^3=-0.15~{\rm GeV}^3$, $\rho_D^3=0.2~{\rm GeV}^3$. For the other input parameters, we employ $m_b^{\rm kin}(1~{\rm GeV})=4.563~{\rm GeV}$, $\overline{m}_c(3~{\rm GeV})=0.989~{\rm GeV}$, $\alpha_s(m_b^{\rm kin})=0.2182$.

3.6 Numerical results for lepton energy and hadronic mass moments

The accuracy of our moment predictions is summarized in Tab. 1. As the NNLO corrections to the q^2 moments are known analytically [47], and a detail discussion was given recently in that reference, we do not discuss these here in detail. However, since our implementation for the lepton energy and hadronic mass moments depends on our fit for the non-BLM contributions, it is interesting to give the relative contributions to the moments. These contributions are given in Tab. 2, where we use $E_{\rm cut}=1~{\rm GeV}$. For the HQE parameters, we use $\mu_{\pi}^2=0.4~{\rm GeV}^2$, $\mu_G^2=0.35~{\rm GeV}^2$, $\rho_{LS}^3=-0.15~{\rm GeV}^3$, $\rho_D^3=0.2~{\rm GeV}^3$. For the other input parameters, we employ $m_b^{\rm kin}(1~{\rm GeV})=4.563~{\rm GeV}$, $\overline{m}_c(3~{\rm GeV})=0.989~{\rm GeV}$, $\alpha_s(m_b^{\rm kin})=0.2182$.

In Tab. 2, we also explicitly give the effect of including the charm decoupling in the kinetic scheme conversion. This contribution is labelled as Δ_c . Our results are in good agreement with Ref. [61] up to the non-BLM corrections. The Δ_c effects were not included in that reference. Currently, these contributions are automatically included at NNLO in kolya, and they cannot be separated from other contributions.

We observe that Δ_c has a 0.1% effect and is of similar in size but opposite in sign as the non-BLM correction. Overall, the α_s^2 contributions have a 0.05% effect.

We provide additional numerical examples and comparisons with literature in the Git-Lab repository, in the Jupyter notebook example-reproduce_literature.ipynb.

4 Extension to physics beyond the SM

In kolya, we also implement NP effects in $b \to cl\bar{\nu}_l$ decays following [12]. In order to parametrize effects beyond the SM, we use the weak effective theory (WET), an effective field theory valid below the EW scale written as an expansion in powers of the inverse electroweak scale $G_F = 1/(\sqrt{2}v^2)$. The effective Hamiltonian relevant for $B \to X_c l\bar{\nu}_l$ is

given by

$$\mathcal{H}_{\text{eff}} = \frac{4G_F V_{cb}}{\sqrt{2}} \left[(1 + C_{V_L}) \,\mathcal{O}_{V_L} + \sum_{i=V_R, S_L, S_R, T} C_i \,\mathcal{O}_i \right], \tag{42}$$

where we consider only the dimension-six operators which contribute to the differential rate at tree level:

$$\mathcal{O}_{V_{L(R)}} = \left(\bar{c}\gamma_{\mu}P_{L(R)}b\right)\left(\bar{\ell}\gamma^{\mu}P_{L}\nu_{\ell}\right) ,$$

$$\mathcal{O}_{S_{L(R)}} = \left(\bar{c}P_{L(R)}b\right)\left(\bar{\ell}P_{L}\nu_{\ell}\right) ,$$

$$\mathcal{O}_{T} = \left(\bar{c}\,\sigma_{\mu\nu}P_{L}b\right)\left(\bar{\ell}\,\sigma^{\mu\nu}P_{L}\nu_{\ell}\right) .$$
(43)

We define $P_{L(R)} = 1/2 (1 \mp \gamma_5)$ and $\sigma^{\mu\nu} = \frac{i}{2} [\gamma^{\mu}, \gamma^{\nu}]$. In the SM, only O_{V_L} contributes. We have written out this contribution explicitly, such that all Wilson coefficients C_i are zero in the SM. We assume that all C_i are real-valued. We do not consider interactions with right-handed neutrinos.

If one would like to parametrize NP effects in the SMEFT framework [66], there would be an additional expansion in powers of $1/\Lambda$, where Λ corresponds to the NP scale above the EW scale. The tree-level matching of SMEFT operators onto the effective Hamiltonian can be obtained from Ref. [67]. In the WET, the expansion parameter is 1/v, therefore from the SMEFT point of view the Wilson coefficients in (42) would be further suppressed by the small ratio $(v/\Lambda)^2$.

The NP contributions to the differential rate of $B \to X_c l \bar{\nu}_l$ have been presented in Ref. [12]. The NP effects for the moments defined in (3), obtained from the integration of the triple differential rate, can be written schematically in the following way

$$\langle O \rangle = \xi_{\text{SM}} + |C_{V_R}|^2 \xi_{\text{NP}}^{\langle V_R, V_R \rangle} + |C_{S_L}|^2 \xi_{\text{NP}}^{\langle S_L, S_L \rangle} + |C_{S_R}|^2 \xi_{\text{NP}}^{\langle S_R, S_R \rangle} + |C_T|^2 \xi_{\text{NP}}^{\langle T, T \rangle}$$

$$+ \operatorname{Re}((C_{V_L} - 1)C_{V_R}^*) \xi_{\text{NP}}^{\langle V_L, V_R \rangle} + \operatorname{Re}(C_{S_L}C_{S_R}^*) \xi_{\text{NP}}^{\langle S_L, S_R \rangle} + \operatorname{Re}(C_{S_L}C_T^*) \xi_{\text{NP}}^{\langle S_L, T \rangle}$$

$$+ \operatorname{Re}(C_{S_R}C_T^*) \xi_{\text{NP}}^{\langle S_R, T \rangle} .$$

$$(44)$$

The coefficients denoted by ξ correspond to the various interference terms between different effective operators. They depend on the bottom and charm quark masses, the HQE parameters and the lepton energy cut or the q^2 cut. Ref. [12] provides results for the power corrections at tree level up to $O(1/m_b^3)$ and NLO perturbative corrections to leading order in the power expansion $(1/m_b^0)$. In (44), we assume that the NP Wilson coefficients C_i are much smaller than unity therefore, when calculating the moments in (3) we expand up to quadratic NP couplings. We note that the contribution $C_{V_L}\xi_{NP}^{(V_L)}$ drops out for such normalized moments while for the branching ratio, it leads to a rescaling of V_{cb} .

The first term ξ_{SM} in (44) corresponds to the SM prediction, whose implementation has been described in the previous sections. The additional NP contributions generated by the effective operators have been implemented in kolya including power corrections up

to $1/m_b^3$ and NLO perturbative QCD corrections at partonic level. The implementation closely follows the methods described for the SM case. Namely, we implement the tree-level contributions in an exact form, while for the QCD corrections, we generate interpolation grids for their fast evaluation. The NP contributions to the moments are implemented in Q2moments_NP.py, Elmoments_NP.py and MXmoments_NP.py. The NP extension of the total rate is in TotalRate_NP.py.

5 Usage of the library

5.1 Installation

The software kolya requires Python version 3.6 or above and runs on Linux and Mac. The code is released under the GNU GPL v3 license. To download the package, clone the master branch of the Git repository via

```
$: git clone https://gitlab.com/vcb-inclusive/kolya.git
```

Afterwards, change the directory to the kolya directory and install it with pip3 in the following way:

```
$: cd kolya
$: pip3 install .
```

The dependencies will be automatically downloaded and installed during the setup. To get started, just import the package into a Python shell or a Jupyter notebook:

```
>>> import kolya
```

Note that the first time kolya is loaded, several functions are translated from Python to optimized machine code by Numba and cached. This stage may take from several seconds up to a few minutes.

5.2 Parameter classes

The library contains classes to store various real-valued variables. One class is dedicated to the physical parameters like heavy quark masses and the strong coupling constant, one for the HQE parameters, and one for the Wilson coefficients in the NP extension. Dimensionful quantities, like the quark masses, are given in units of GeV. The values of the physical parameters are stored in an object of parameters.physical_parameters class

```
>>> par = kolya.parameters.physical_parameters()
```

The new object par contains information about M_B , $m_b^{\text{kin}}(\mu_{\text{kin}})$, $\overline{m}_c(\mu_c)$ and $\alpha_s^{(4)}(\mu_s)$. The bottom quark mass in the kinetic scheme and its scale are given in the variables mbkin and scale_mbkin, while the charm mass in the $\overline{\text{MS}}$ at renormalization scale μ_c corresponds to

the variables mcMS and scale_mcMS. The strong coupling constant $\alpha_s^{(4)}$ and its renormalization scale scale μ_s correspond to the class variables alphas and scale_alphas. The class initializes also the renormalization scales of the HQE parameters μ_G^2 , ρ_D^3 and ρ_{LS}^3 through the variables scale_muG, scale_rhoD and scale_rhoLS. At present, these are set equal to $m_b^{\rm kin}(\mu_{\rm kin})$.

The values stored in the object par can be shown with the show method:

```
>>> par.show() bottom mass: mbkin(1.0 \text{ GeV}) = 4.563 \text{ GeV} charm mass: mcMS(3.0 \text{ GeV}) = 0.989 \text{ GeV} coupling constant: alpha_s(4.563 \text{ GeV}) = 0.2182
```

where the current default values are based on the latest version of the FLAG 21 review [68] as of February 2024 [69]. Values different from the default ones can be set during initialization. For instance, we can initialize the $\overline{m}_c(2 \text{ GeV}) = 1.094 \text{ GeV}$ as follows:

```
>>> par = kolya.parameters.physical_parameters(mcMS=1.094,scale_mcMS=2.0)
>>> par.show()
bottom mass:     mbkin( 1.0 GeV) = 4.563 GeV
charm mass:     mcMS( 2.0 GeV) = 1.094 GeV
coupling constant: alpha_s( 4.563 GeV) = 0.2182
```

The example above shows that during initialization, the values of mcMS and scale_mcMS must be set consistently. The following command

```
>>> par = kolya.parameters.physical_parameters(mcMS=1.094)
```

would initialize the charm mass to the (unphysical) value $\overline{m}_c(3\,\text{GeV}) = 1.094\,\text{GeV}$. In order to set the quark masses at scales different from the default ones in a consistent way, we include the method FLAG2024. For instance, we set the quark masses at a scale $\mu_{\text{kin}} = \mu_c = 2\,\text{GeV}$ in the following way:

```
>>> par = kolya.parameters.physical_parameters()
>>> par.FLAG2024(scale_mcMS=2.0, scale_mbkin=2.0)
>>> par.show()
bottom mass:          mbkin( 2.0 GeV) = 4.295730717092438 GeV
charm mass:          mcMS( 2.0 GeV) = 1.0940623249384822 GeV
coupling constant: alpha_s( 4.563 GeV) = 0.21815198098622618
```

Internally, the bottom and quark masses are recalculated using CRunDec [70–72] using the default values from Ref. [69]. The scale of the strong coupling constant can be modified in a similar way:

Also in this case, we internally use CRunDec to evaluate $\alpha_s^{(4)}(3 \,\text{GeV})$.

The values of the HQE parameters in the historical basis (sometimes referred to as the "perp" basis in literature) are stored into an object of the class parameters.HQE_parameters.By default, their values are set to zero unless explicitly initialized:

```
>>> hqe = kolya.parameters.HQE_parameters(
               muG = 0.306,
               rhoD = 0.185,
               rhoLS = -0.13,
               mupi = 0.477)
>>> hqe.show()
mupi = 0.477
                    GeV<sup>2</sup>
muG
           0.306
                    GeV<sup>2</sup>
           0.185
                    GeV<sup>3</sup>
rhoLS =
          -0.13
                    GeV<sup>3</sup>
```

where the values up to $1/m_b^3$ can be visualized all at once with $\mathtt{show}()$. Since there are several operators at order $1/m_b^4$ and $1/m_b^5$, we do not print them by default, however, their values can be inspected with the additional option $\mathtt{show}(\mathtt{flagmb4=1})$ and $\mathtt{show}(\mathtt{flagmb5=1})$. We introduce the class $\mathtt{parameters.HQE_parameters_RPI}$ for the HQE parameters in the RPI basis:

The parameter classes LSSA_HQE_parameters and LSSA_HQE_parameters_RPI contain numerical values for the HQE parameters in the historical and RPI basis, respectively, obtained using the "lowest-lying state saturation ansatz" (LLSA). The LLSA approximates the higher-order HQE parameters by expressing them through the bottom quark mass m_b , the HQE parameters $(\mu_{\pi}^2)^{\perp}$ and $(\mu_G^2)^{\perp}$, and the excitation energies $\epsilon_{1/2}$ and $\epsilon_{3/2}$. For further details on the LLSA, we refer to Ref. [73], and for the expressions of the higher-order HQE parameters and their LLSA values, we refer to Refs. [28, 32].

The Wilson coefficients of the effective dimension-six operators defined in (42) are initialized via the class parameters. WCoefficients and their values are inspected with the method show()

```
>>> wc = kolya.parameters.WCoefficients(SL=0.1,SR=-0.05)
>>> wc.show()
C_{V_L} = 0
C_{V_R} = 0
C_{S_L} = 0.1
C_{S_R} = -0.05
C_{T} = 0
```

The Wilson coefficients C_{V_L} , C_{V_R} , C_{S_L} , C_{S_R} and C_T are denoted by VL, VR, SL, SR and T respectively. By default, they are initialized to zero.

5.3 Moment predictions

We implemented the first four centralized moments of the q^2 spectrum and the first three moments of E_l and M_X^2 . To evaluate them, we first need to initialize three objects for the physical parameters, the HQE parameters, and the Wilson coefficients:

The prediction for the q^2 moments receive as inputs the value of $q_{\rm cut}^2$ expressed in GeV², and the three objects par, hqe and wc. The value of each q^2 moment is obtained with the functions Q2moments.moment_n_KIN_MS(q2cut, par, hqe, wc), where n= 1, 2, 3, 4. For example, to evaluate $q_1(q_{\rm cut}^2)$ for $q_{\rm cut}^2=8.0~{\rm GeV}^2$, type

```
>>> q2cut = 8.0  #GeV^2
>>> kolya.Q2moments.moment_1_KIN_MS(q2cut, par, hqe, wc)
8.996406491856465 #GeV^2
```

The result is provided in the respective powers of GeV^{2n} . The suffix KIN and MS refers to the scheme for bottom (kinetic) and charm ($\overline{\text{MS}}$) masses. By default, the evaluation considers power corrections up to $1/m_b^3$. The corrections of order $1/m_b^4$ and $1/m_b^5$ can be included by setting the optional arguments flagmb4=1 and flagmb5=1. For instance, we can set the higher-order HQE parameters $m_1 = 0.1 \,\text{GeV}^4$ and $r_1 = 0.1 \,\text{GeV}^5$ and compare the predictions up to order $1/m_b^5$ in the following way:

```
>>> hqe.m1=0.1 #GeV^4
>>> hqe.r1=0.1 #GeV^5
>>> kolya.Q2moments.moment_1_KIN_MS(q2cut, par, hqe, wc)
8.996406491856465
>>> kolya.Q2moments.moment_1_KIN_MS(q2cut, par, hqe, wc, flagmb4=1)
8.966491804532277
>>> kolya.Q2moments.moment_1_KIN_MS(q2cut, par, hqe, wc, flagmb4=1, flagmb5=1)
8.75600814954277
```

Setting flagmb4=0 and flagmb5=0 eliminates all terms of order $1/m_b^4$ and $1/m_b^5$, respectively. We note that at these orders also mixing terms proportional to $\mu_G^2 \mu_\pi^2$ or $\mu_G^2 \rho_D^3$ enter which can only be excluded by putting these two flags to zero. Therefore, putting these

flags to zero does not have the same effect as simply setting all the $1/m_b^4$ and $1/m_b^5$ HQE element to zero.

Concerning the perturbative corrections, these are all included by default in the numerical evaluation (see Tab. 1 for the current orders in α_s implemented). For cross-checks with the literature and the study of their impact, the NNLO corrections can be switched off via the optional argument flag_includeNNLO=0 (the default is flag_includeNNLO=1). Also, the NLO corrections to the power-suppressed terms can be excluded with flag_includeNLOpw=0.

Moreover, the option flag_DEBUG=1 will print a report of the various contributions coming from the higher-order QCD corrections:

```
>>> kolya.Q2moments.moment_1_KIN_MS(8.0, par, hqe, wc, flag_DEBUG=1)
Q2moment n. 1 LO = 9.148659808170105
Q2moment n. 1 NLO = api * -1.319532010835962
Q2moment n. 1 NNLO = api^2 * -9.616956902561078
Q2moment n. 1 NLO pw = api * -0.7873907726673756
Q2moment n. 1 NNLO from NLO pw = api^2 * 8.39048437244325
```

The contributions denoted by NLO and NNLO are the coefficients in front of $\alpha_s(\mu_s)/\pi$ and $(\alpha_s(\mu_s)/\pi)^2$ to leading order in $1/m_b$. The term NLO pw corresponds to the overall NLO correction in the terms of order $1/m_b^2$ and $1/m_b^3$. In the kinetic scheme, the inclusion of the NLO corrections to the power-suppressed terms induces also an additional $O(\alpha_s^2)$ contribution to leading order in $1/m_b$, which is reported in the last line.

The predictions for the E_l and M_X^2 moments follow a similar syntax. The first argument passed to the function corresponds to the value of the cut E_{cut} in units of GeV. For instance, the first moments of E_l and M_X^2 for $E_{\text{cut}} = 1.0$ GeV are evaluated as follows:

Higher moments ℓ_2, ℓ_3, h_2, h_3 are computed in a similar way by replacing moment_1 with moment_2 or moment_3. The result for $\ell_n(E_{\text{cut}})$ is in GeV^n , while for $h_n(E_{\text{cut}})$ the result is in GeV^{2n} .

By default, the moments are calculated using the HQE parameters as defined in the historical basis. The q^2 moments and the total rate can also be calculated using the RPI basis adopted in Ref. [9]. The predictions in the RPI basis are obtained by passing the

optional argument flag_basisPERP=0. In this case, the HQE parameters must be passed to the function through an object of the class HQE_parameters_RPI:

The RPI basis is supported only for q_n moments and the total rate since reparametrization invariance reduces the number of HQE parameters only for them. Similar to before, the $1/m_b^4$ and $1/m_b^5$ corrections in the RPI basis can be included by using the optional arguments flagmb4=1 and flagmb5=1. For the ℓ_n and h_n , we stick to the historical basis.

5.4 Branching ratio prediction

To obtain the branching ratio $Br(B \to X_c l \bar{\nu}_l)$ or the total semileptonic width $\Gamma_{\rm sl}$, three objects for the physical parameters, the HQE parameters, and the Wilson coefficients must be initialized, as discussed for the moments in the previous subsection. For the total rate $\Gamma_{\rm sl}$ defined in (2), type

```
>>> Vcb = 42.2e-3
>>> kolya.TotalRate.TotalRate_KIN_MS(Vcb, par, hqe, wc)
4.4016320941077224e-14 #GeV
```

where the first argument is the value of $|V_{cb}|$ and the result is expressed in GeV. For the evaluation of the branching ratio, we use

```
>>> Vcb = 42.2e-3
>>> kolya.TotalRate.BranchingRatio_KIN_MS(Vcb, par, hqe, wc)
0.10555834162102022
```

We obtain the branching ratio by dividing $\Gamma_{\rm sl}$ by the average lifetime of the B^{\pm} and B_0 mesons.

The partial width $\Delta\Gamma_{\rm sl}(E_{\rm cut})$ with cut on E_l is obtained with

```
>>> Vcb = 42.2e-3
>>> Elcut = 1.0 # GeV
>>> kolya.DeltaBR.DeltaRate_KIN_MS(Vcb, Elcut, par, hqe, wc)
3.469103884950744e-14 #GeV
```

where the first argument is the value of $|V_{cb}|$, the second argument the value of E_{cut} in GeV. The result is reported in GeV. The corresponding value for the branching ratio is given by

```
>>> Vcb = 42.2e-3
>>> Elcut = 1.0 # GeV
>>> kolya.DeltaBR.DeltaBR_KIN_MS(Vcb, Elcut, par, hqe, wc)
0.08319478892764474
```

The function that predicts the branching ratio allows optional arguments flagmb4=1 and flagmb5=1 to include the power corrections of order $1/m_b^4$ and $1/m_b^5$. The predictions in the RPI basis are obtained by passing the optional argument $flag_basisPERP=0$. For cross-checks with the literature and the study of the impact of QCD corrections, the NNLO and N3LO corrections to the total rate can be switched off via the optional arguments $flag_includeNNLO=0$ and $flag_includeN3LO=0$ (by default, all these corrections are included). Moreover, the effects arising from the NLO corrections to the power-suppressed terms can be excluded with $flag_includeNLOpw=0$.

6 Outlook & Conclusion

In this document, we have presented the first version of the open-source library kolya, corresponding to the release 1.0. In this release, we have implemented the predictions in the HQE for the total rate and the moments of q^2 , E_l and M_X^2 . Currently, this is sufficient for comparison with published experimental results by B factories. We included all higher order corrections in α_s and $1/m_b$ which are available at this specific point in time and are summarized in Tab. 1.

On the GitLab repository, we provide, additionally, interactive tutorials running as a Jupyter notebook and validation notebooks which demonstrate how the library can reproduce the results available in the literature. The library is open source, so code contributions and improvements are very welcome. In particular, new higher-order corrections can be implemented like

- QED effects calculated in Ref. [74],
- exact results for the NNLO corrections to E_l and M_X^2 moments with a lower cut E_{cut} ,
- renormalization group evolution of the HQE parameters to NLO,
- the NLO corrections in the coefficients of ρ_D^3 and ρ_{LS}^3 for the E_l and M_X^2 moments.

Additional observables can play an important role in better improving the extraction of the HQE parameters or have an important role in testing the SM. These new observables may include

- forward-backward asymmetries A_{FB} and q^2 , E_l and M_X^2 moments for forward and backward events [75, 76],
- the ratio $R_X = \Gamma_{B \to X_c \tau \bar{\nu}_\tau} / \Gamma_{B \to X_c l \bar{\nu}_l}$,
- the lifetime of B mesons within the HQE,

• predictions for the decay into charmless final states $B \to X_u l \bar{\nu}_l$.

Finally, kolya could also be extended to include predictions for inclusive D decays discussed in detail in [77].

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A Definition of the HQE elements

Here, we define the HQE parameters both in the historical and the RPI basis up to $1/m_b^5$. The conversions between these two bases can be found in Ref. [28].

A.1 Historical basis

The HQE matrix elements in the historical basis, denoted by " \perp ", are defined through the spacial covariant derivatives $iD_{\mu}^{\perp} = g_{\mu\nu}^{\perp}iD^{\nu}$, where

$$g_{\mu\nu}^{\perp} = g_{\mu\nu} - v_{\mu}v_{\nu} , \qquad (45)$$

as in Refs. [27, 79]. We will employ here the notation $\langle \bar{b}_v \dots b_v \rangle \equiv \langle B(v) | \bar{b}_v \dots \bar{b}_v | B(v) \rangle$. At $1/m_h^2$, we have

$$2m_B(\mu_\pi^2)^{\perp} = -\langle \bar{b}_v (iD^{\rho}) (iD^{\sigma}) b_v \rangle g_{\rho\sigma}^{\perp} ,$$

$$2m_B(\mu_G^2)^{\perp} = \frac{1}{2} \langle \bar{b}_v [(iD^{\rho}), (iD^{\sigma})] (-i\sigma^{\alpha\beta}) b_v \rangle g_{\rho\alpha}^{\perp} g_{\sigma\beta}^{\perp} ,$$
(46)

where $\gamma^{\mu}\gamma^{\nu} = g^{\mu\nu} + (-i\sigma^{\mu\nu})$. At $1/m_b^3$, we have

$$2m_B(\rho_D^3)^{\perp} = \frac{1}{2} \langle \bar{b}_v \left[(iD^{\rho}), \left[(iD^{\sigma}), (iD^{\lambda}) \right] \right] b_v \rangle g_{\rho\lambda}^{\perp} v_{\sigma} ,$$

$$2m_B(\rho_{LS}^3)^{\perp} = \frac{1}{2} \langle \bar{b}_v \left\{ (iD^{\rho}), \left[(iD^{\sigma}), (iD^{\lambda}) \right] \right\} (-i\sigma^{\alpha\beta}) b_v \rangle g_{\rho\alpha}^{\perp} g_{\lambda\beta}^{\perp} v_{\sigma} . \tag{47}$$

The nine HQE parameters at $1/m_b^4$ were first introduced in Ref. [27]. We list them here:

$$2m_{B}m_{1} = \langle \bar{b}_{v} \left(iD^{\rho} \right) \left(iD^{\sigma} \right) \left(iD^{\delta} \right) b_{v} \rangle \frac{1}{3} \left(g_{\rho\sigma}^{\perp} g_{\lambda\delta}^{\perp} + g_{\rho\lambda}^{\perp} g_{\sigma\delta}^{\perp} + g_{\rho\delta}^{\perp} g_{\sigma\lambda}^{\perp} \right) ,$$

$$2m_{B}m_{2} = \langle \bar{b}_{v} \left[(iD^{\rho}), (iD^{\sigma}) \right] \left[(iD^{\lambda}), (iD^{\delta}) \right] b_{v} \rangle g_{\rho\delta}^{\perp} v_{\sigma} v_{\lambda} ,$$

$$2m_{B}m_{3} = \langle \bar{b}_{v} \left[(iD^{\rho}), (iD^{\sigma}) \right] \left[(iD^{\lambda}), (iD^{\delta}) \right] b_{v} \rangle g_{\rho\lambda}^{\perp} g_{\sigma\delta}^{\perp} ,$$

$$2m_{B}m_{4} = \langle \bar{b}_{v} \left\{ (iD^{\rho}), \left[(iD^{\sigma}), \left[(iD^{\lambda}), (iD^{\delta}) \right] \right] \right\} b_{v} \rangle g_{\sigma\lambda}^{\perp} g_{\rho\delta}^{\perp} ,$$

$$2m_{B}m_{5} = \langle \bar{b}_{v} \left[(iD^{\rho}), (iD^{\sigma}) \right] \left[(iD^{\lambda}), (iD^{\delta}) \right] \left(-i\sigma^{\alpha\beta} \right) b_{v} \rangle g_{\alpha\rho}^{\perp} g_{\beta\delta}^{\perp} v_{\sigma} v_{\lambda} ,$$

$$2m_{B}m_{6} = \langle \bar{b}_{v} \left[(iD^{\rho}), (iD^{\sigma}) \right] \left[(iD^{\lambda}), (iD^{\delta}) \right] \left(-i\sigma^{\alpha\beta} \right) b_{v} \rangle g_{\alpha\sigma}^{\perp} g_{\beta\lambda}^{\perp} g_{\rho\delta}^{\perp} ,$$

$$2m_{B}m_{7} = \langle \bar{b}_{v} \left\{ \left\{ (iD^{\rho}), (iD^{\sigma}) \right\}, \left[(iD^{\lambda}), (iD^{\delta}) \right] \right\} \left(-i\sigma^{\alpha\beta} \right) b_{v} \rangle g_{\rho\sigma}^{\perp} g_{\alpha\lambda}^{\perp} g_{\beta\delta}^{\perp} ,$$

$$2m_{B}m_{8} = \langle \bar{b}_{v} \left\{ \left\{ (iD^{\rho}), (iD^{\sigma}) \right\}, \left[(iD^{\lambda}), (iD^{\delta}) \right] \right\} \left(-i\sigma^{\alpha\beta} \right) b_{v} \rangle g_{\rho\sigma}^{\perp} g_{\alpha\lambda}^{\perp} g_{\beta\delta}^{\perp} ,$$

$$2m_{B}m_{9} = \langle \bar{b}_{v} \left\{ \left\{ (iD^{\rho}), \left[(iD^{\sigma}), \left[(iD^{\lambda}), (iD^{\delta}) \right] \right\} \right\} \left(-i\sigma^{\alpha\beta} \right) b_{v} \rangle g_{\rho\sigma}^{\perp} g_{\alpha\lambda}^{\perp} g_{\sigma\delta}^{\perp} . \tag{48}$$

Finally, eighteen more parameters are present at $1/m_b^5$, as defined in Ref. [27]:

$$2m_{B}r_{1} = \langle \bar{b}_{v} (iD_{\mu}) (ivD)^{3} (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{2} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD^{\mu}) (iD)^{2} b_{v} \rangle, \\
2m_{B}r_{3} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD^{\mu}) (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{4} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD^{\mu}) (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{5} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD)^{2} (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{5} = \langle \bar{b}_{v} (iD^{\mu}) (iD^{\mu}) (iD^{\mu}) (iD^{\mu}) (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{6} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\nu}) (ivD) (iD^{\mu}) (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{7} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\nu}) (ivD) (iD^{\mu}) (iD^{\mu}) b_{v} \rangle, \\
2m_{B}r_{8} = \langle \bar{b}_{v} (iD_{\mu}) (ivD)^{3} (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{9} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\beta}) (iD)^{2} (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{10} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\mu}) (iD_{\mu}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{11} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\mu}) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{12} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\mu}) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{13} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\beta}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{14} = \langle \bar{b}_{v} (iD_{\mu}) (ivD) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{15} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (ivD) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{15} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (ivD) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{15} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (ivD) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{15} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\beta}) (iD^{\mu}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{17} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\beta}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{17} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle, \\
2m_{B}r_{17} = \langle \bar{b}_{v} (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\mu}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle.$$
(49)

A.2 RPI basis

The RPI HQE matrix elements up to $1/m_b^4$ have been determined in Ref. [31]. We list them here:

$$2m_{B}\mu_{\pi}^{2} = -\langle \bar{b}_{v} (iD)^{2} b_{v} \rangle ,$$

$$2m_{B}\mu_{G}^{2} = \langle \bar{b}_{v} (iD_{\alpha}) (iD_{\beta}) (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}\tilde{\rho}_{D}^{3} = \frac{1}{2} \langle \bar{b}_{v} \left[(iD_{\mu}), \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD^{\mu}) \right] \right] b_{v} \rangle ,$$

$$2m_{B}r_{G}^{4} = \langle \bar{b}_{v} \left[(iD_{\mu}), (iD_{\nu}) \right] \left[(iD^{\mu}), (iD^{\nu}) \right] b_{v} \rangle ,$$

$$2m_{B}\tilde{r}_{E}^{4} = \langle \bar{b}_{v} \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD_{\mu}) \right] \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD^{\mu}) \right] b_{v} \rangle ,$$

$$2m_{B}s_{B}^{4} = \langle \bar{b}_{v} \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD_{\alpha}) \right] \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}\tilde{s}_{E}^{4} = \langle \bar{b}_{v} \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD_{\alpha}) \right] \left[\left((ivD) + \frac{1}{2m_{b}} (iD)^{2} \right), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}s_{qB}^{4} = \langle \bar{b}_{v} \left[\left((iD_{\mu}), \left[(iD^{\mu}), \left[(iD_{\alpha}), (iD_{\beta}) \right] \right] \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$(50)$$

We note that we choose to use μ_{π}^2 instead of μ_3 as in Ref. [31], in order to avoid factors of $1/\mu_3$ in the centralized moments. Furthermore, we note that $\tilde{\rho}_D^3$, \tilde{r}_E^4 , and \tilde{s}_E^4 contain their so-called RPI-completion terms as described in Refs. [28, 31]. In Ref. [28], the RPI matrix

elements at $1/m_b^5$ have been determined to be:

$$2m_{B}X_{1}^{5} = \langle \bar{b}_{v} \left[(ivD), \left[(ivD), (iD_{\mu}) \right] \right] \left[(ivD), (iD^{\mu}) \right] b_{v} \rangle ,$$

$$2m_{B}X_{2}^{5} = \langle \bar{b}_{v} \left[(ivD), \left[(iD_{\mu}), (iD_{\nu}) \right] \right] \left[(iD^{\mu}), (iD^{\nu}) \right] b_{v} \rangle ,$$

$$2m_{B}X_{3}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(ivD), (iD_{\nu}) \right] \left[(iD^{\mu}), (iD^{\nu}) \right] \right] b_{v} \rangle ,$$

$$2m_{B}X_{4}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(iD_{\nu}), \left[(ivD), (iD^{\mu}), (iD^{\nu}) \right] \right] \right] b_{v} \rangle ,$$

$$2m_{B}X_{5}^{5} = \langle \bar{b}_{v} \left[(ivD), \left[(ivD), (iD_{\alpha}) \right] \right] \left[(ivD), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}X_{6}^{5} = \langle \bar{b}_{v} \left[(ivD), \left[(iD_{\mu}), (iD_{\alpha}) \right] \right] \left[(iD^{\mu}), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}X_{7}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(ivD), (iD_{\alpha}) \right] \right] \left[(iD^{\mu}), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}X_{8}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(ivD), (iD^{\mu}) \right] \right] \left[(iD_{\alpha}), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}X_{9}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(ivD), (iD^{\mu}) \right] \right] \left[(iD_{\alpha}), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle ,$$

$$2m_{B}X_{10}^{5} = \langle \bar{b}_{v} \left[(iD_{\mu}), \left[(ivD), (iD^{\mu}) \right] \right] \left[(iD_{\alpha}), (iD_{\beta}) \right] (-i\sigma^{\alpha\beta}) b_{v} \rangle .$$

$$(51)$$

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